








What have we learned from space missions to asteroids?

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Asteroid (Type)	Gaspra (S)	Mathilde (C)	Ida (S)	Eros (S)	Itokawa (S)	Steins (E)	Lutetia (M? C?)
							
Diameter	12 km	53 km	31 km	17 km	0.35 km	6.7 x 5.9 x 4.3 km	121 x 112 x 87 km
Period	7.09 hr	17.406 d	4.634 hr	5.267 hr	12.132 hr	6.047 hr	8.168 hr
Age	200 My	2-4.5 Gy	1 Gy	2 Gy	1-100 My	100-150 My	0.07-3,5 Gy
Density	2.7g/cm ³ (b)	1.3 g/cm ³ (a)	2.6 g/cm ³ (b)	2.67 g/cm ³ (b)	1.95 g/cm ³ (b)	? (c)	3,5-3.9 g/cm ³
Porosity	?	55 – 63 %	18 – 24 %	16 – 21 %	39 – 43 %	?	?
Meteorite	ordinary chondrite	carbonaceous chondrite	ordinary chondrite	ordinary chondrite	ordinary chondrite	aubrite	? CR/CO/CV or Enstatite chondrite?
Objective	Fly-By Galileo (1991) Res=54m/px	Fly-by NEAR (1997) Res=180m/px	Fly-by Galileo (1993) Res=25m/px	1 year-RD NEAR (2000) Res=cm/px	Hovering Hayabusa (2005) Res<1cm/px	Fly-by Rosetta (2008) Res<80 m/px	Fly-by Rosetta (2010) Res >60 m/px
	-First asteroid with young age (200 Myr) -Absence of large craters	-First asteroid with low density - Large craters (5 with D> 5 km) suggest porous bodies have much higher impact strength than expected	- First discovery of a satellite (Dactyl) - Age estimate (1 Byr) - First estimate of density of S-type - First constraints on mechanical properties	- Larger amount of boulders than expected - Lack of very small craters - First evidence of thick regolith	- First evidence of rubble-pile structure - First S-type with low bulk density - Large boulders - Lack of small craters (<10 m) requires unknown process	-- First chunk of a highly differentiated object --First visit to a body shaped by the YORP effect?	--Larger, older explored asteroid --high density -- heterogeneity -- Very large craters (D>40 km) -- Landslides --Fields of large boulders (> 60 m)

(a) Average densities of meteorites for C type asteroids: 2.9 – 3.5 g/cm³

(b) Average densities of meteorites for S type asteroids: 3.19 – 3.40 g/cm³

(c) Average densities of aubrites 2.97 – 3.27 g/cm³